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Atmospheric mercury over the NE Pacific during spring 2002: Gradients, residence time, upper troposphere lower stratosphere loss, and long-range transport

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1 Gaseous elemental mercury (GEM) was measured over the North Pacific and coastal California, Oregon and Washington in spring 2002 during the Intercontinental Transport and Chemical Transformation Experiment–2002 (ITCT2K2). Observations were made from the NOAA WP3 on 11 flights which spanned an altitude range of 0–8 km and latitudes between 35 and 55°N. GEM was determined using a Tekran 2537A mercury instrument modified for airborne use to collect and analyze samples every 2.5 min under variable ambient temperature and pressure conditions. GEM concentrations were in general at the low end of previous measurements made over continents and in fresh continental outflow over the ocean. They showed (1) a significant vertical gradient in GEM between the lower sample altitudes and the upper altitudes, with median concentrations decreasing from 1.2 to 0.5 ng m⁻³; (2) GEM was depleted in upper troposphere lower stratospheric air and indicated this region is a sink for GEM; (3) GEM concentrations were ~1/2 at the surface to ~1/4 at midaltitudes of those observed immediately downwind of Asia during ACE-Asia, however; (4) highest GEM concentrations (~4 ng m⁻³) were found in plume layers at concentrations comparable to those observed off the Asian coast in ACE-Asia; and (5) the observed variance in GEM suggested a lifetime on the order of 100 days, which was shorter than the previously accepted nominal lifetime of 1 year.


1. Introduction

[2] Mercury, Hg, is the only significant metallic atmospheric vapor and is typically present in air as gaseous elemental mercury, GEM, with a typical concentration between 1 to 2 ng m⁻³. GEM concentrations are usually reported in nanograms per standard cubic meter where standard temperature and pressure are 0°C and 1013.25 hPa, respectively. GEM has an oxidation state of zero and accounts for 98% of the tropospheric mercury burden. Particulate mercury (pHg) and gaseous ionic or reactive mercury (RGM) comprise the bulk of the remaining 2%. The residual atmospheric mercury is composed of the more highly toxic organic Hg forms like methylmercury. Regardless of form, mercury is a global-scale pollutant.

[3] GEM is released into the atmosphere from a multitude of sources, for example, fugitive emissions from industrial processes, coal combustion, waste incineration, biomass fires, and volcanism. GEM is oxidized to RGM by hydroxyl, halogen, and nitrogen trioxide radicals and to a lesser extent by ozone [e.g., Pal and Ariya, 2004]. Hg in its more oxidized forms (Hg IV, e.g., HgO, HgSO₄, HgCl₂) can be further converted into pHg. These compounds are thought to be relatively short-lived as the expected RGM to particle lifetime is ~1 day and particle lifetimes are of order 10 days [Fitzgerald et al., 1991; Hall, 1995; Schroeder and Munthe, 1998; Bergan and Rodhe, 2001; Nriagu and Becker, 2003].

[4] Elemental and oxidized mercury can be biologically and chemically converted in soils, fresh waters, oceans, and sediments to more highly reactive compounds and to the very toxic organic forms such as methylmercury. The organic forms of Hg are strongly “bioaccumulated” and concentrated up the food chain [Schroeder et al., 1991; Pilson, 1998], with occasionally disastrous results to human populations [Takizawa, 1979].

[5] Mercury tends to be released back into the air after deposition on Earth’s surfaces and mercury’s fluxes between environmental compartments are bidirectional [Mason et al., 1994; Richardson, 2001; Friedli et al., 2003a, 2003b]. Depostied mercury is revolatized by biological methylation, chemical reduction, photolysis and physical processes such as wildfire [Friedli et al., 2001]. Also, it remains unclear whether the global ocean is a net
source or a net sink for atmospheric GEM, but the fluxes between atmosphere and ocean are significant [Fitzgerald, 1986; Temme et al., 2003; Mason and Sheu, 2002; Mason et al., 2001, 2002; Rasmussen, 1994; Pilson, 1998].

[5] Previous observations of mercury in the troposphere and the interpretation of its transport and atmospheric chemistry have lead others to conclude the lifetime of GEM is ~1 year [e.g., Mason et al., 1994; Bergan and Rodhe, 2001] with estimates ranging from 0.5 to 2 years. This lifetime is under challenge. Recent measurements of ambient GEM in the Arctic [Lindberg et al., 2002a], ACE Asia, and forest fires [Lindberg et al., 2002b; Friedli et al., 2004] have suggested GEM is more variable than a 1-year lifetime would predict. Laboratory kinetic studies [e.g., Pal and Ariya, 2004] which considered halogen and NO3 radical chemistry suggest the GEM lifetime should be much shorter, on the order of a few months, depending upon the assumed atmospheric level of hydroxyl, halogen, or nitrogen trioxide radicals. Hedgcock and Pirrone [2004] considered GEM oxidation by Br, O3, and OH and calculated boundary layer residence times as low as 10 days. When only oxidation by OH and O3 were considered, they estimated a GEM lifetime of 133 days. Selin et al. [2007] using the GEOS-CHEM system estimate a GEM global mean annual lifetime significantly shorter than 1 year and were only able to achieve an 8-month lifetime using a model parameter set selected to maximize GEM lifetime. These studies indicated GEM may have a much more dynamic character and call into question the details of current global atmospheric mercury budgets based upon a 1-year GEM lifetime [e.g., Lamborg et al., 2002a].

[6] There are only a handful of observations above the boundary layer with which to evaluate Hg budgets and chemical-transport theory. Tropospheric GEM observations are presented here from the Intercontinental Transport and Chemical Transformation Experiment–2002 (ITCT2K2). They were made aboard the NOAA WP-3 aircraft flown over the eastern North Pacific and along the west coast of the United States in spring 2002. ITCT2K2 was designed to examine the background composition of air before it had been modified by fresh North American emissions and to examine the long-range transport of pollutants from other continents. Observations were made on 11 flights, which spanned an altitude range of 0–8 km and latitudes between 35 and 55°N. The flights went as far south as Los Angeles and as far north as the Northwest tip of Washington and traversed ~50,000 km. Airborne in situ GEM samples were analyzed using a modified Tekran 2537A mercury analyzer. Below, the modifications and testing of the instrument are described first. The ITCT2K2 GEM composition of air coming ashore to North America is presented next and the GEM data are then contrasted with previous observations, in particular with those made off the Asian-Pacific coast in spring 2000 as part of the Asian Pacific Regional Aerosol Characterization Experiment, ACE-Asia [Friedli et al., 2004] and those from Mt. Batchelor, OR [Weiss-Penzias et al., 2003, 2006].

2. Methods

[7] GEM was measured with a Tekran 2537A mercury vapor analyzer (Tekran Inc., Toronto, Canada) modified to eliminate the known pressure dependence of the standard instrument. Figure 1 shows a schematic of the sampling and measurement arrangement used on the aircraft and is described in detail below. Sample air was brought into the aircraft via a rear-facing ¼ inch Teflon tube inlet and plumbed to the instrument. It then passed through a 47 mm, 0.2 μm Teflon filter in a Teflon housing to remove particles. Sample air then passed through one of two gold cartridges where gaseous mercury was collected. The processed air was discharged into a dump line and vented overboard. The sample airflow rate from inlet to dump was regulated by the Tekran’s internal pump. This pump was programmed to draw 1.50 slpm (standard liters per minute, T = 273.15 K, P = 1013.25 hPa). Sample air was replaced by mercury-free zero air for instrument startup and instrument background checks. The zero air was generated by drawing ambient air through a Resisorb activated carbon canister (Tekran Inc.). The Tekran internal pump was used to draw ambient sample air and to generate zero air. Ultrahigh-purity Ar (99.999%, Scott Specialty Gases) was used as the carrier gas for gold cartridge desorption and the desorbed mercury was quantified by cold vapor atomic fluorescence spectroscopy. The Ar pressure in the desorbing gold cartridge and in the fluorescence cell was maintained at 1026 hPa (1.013 atm) using a MKS Series 640 electronic pressure controller, which was plumbed to the outlet of the fluorescence cell. Detector Ar pressure control was a significant modification and the rationale and testing of this modification are explained more fully below.

2.1. Tekran 2537A Operation

[8] The Tekran 2537A instrument concentrated mercury from a metered air sample flow by amalgamating it on a gold cartridge. There are two gold cartridges (“A” and “B”) in parallel flow paths. These were cycled between a collection state and an analysis state to permit continuous sampling every 150 s. In state 1, sample air passed through the cartridge, GEM was collected and the spent air exhausted. In state 2, an isolated Ar carrier stream was passed through the cartridge to displace sample air. After a set period of Ar flow, the cartridge was rapidly heated desorbing the previously collected elemental mercury into the Ar carrier. The thermally desorbed mercury was passed through a fluorescence cell and quantified by cold vapor atomic fluorescence spectroscopy. The instrumental detection limit was about 0.1 ng m⁻³ (GEM concentrations are reported in nanograms per standard cubic meter where standard temperature and pressure are 0°C and 1013.25 hPa, respectively).

[9] The output from the Tekran 2537A instrument was designated as total gaseous mercury (TGM), which includes GEM and the unknown fraction of reactive gaseous mercury (RGM) not scavenged on internal instrument surfaces. RGM is generally present in the atmosphere as a small fraction of GEM [Lindberg and Stratton, 1998]. There are exceptions as Lindberg et al. [2002a, 2002b] observed elevated RGM fractions during Arctic mercury depletion events. A 0.2 μm Teflon filter is placed ahead of the mercury analyzer to remove particulate mercury and potential aerosol contamination from the sample. Thus the GEM measurements were effectively composed of gaseous elemental mercury alone and referred to as GEM. We have
assumed the RGM fraction was small, but if not RGM would constitute an unquantified positive artifact in our GEM data.

The instrument was calibrated premission and postmission using an internal GEM permeation cell and the calibration was verified by manual injection of known amounts of GEM supplied from a Tekran 2505 Gaseous Mercury Calibrator (Tekran Inc., Toronto, Canada). The permeation cell and manual injections agreed to within 1% and 2% for the A and B cartridges, respectively. Calibration factors before the experiment were $6166 \pm 113$ counts/picogram for cartridge A ($n = 13$), $6085 \pm 179$ counts/picogram ($n = 13$) for B and were stable. In the field, permeation cell calibrations and manual calibrations were not possible because of logistical constraints. Hence throughout the field phase of the experiment, the instrument used the last permeation cell calibration as the default, $A = 6351$ cts/pg and $B = 6319$ cts/pg. At the end of the field experiment, the instrument was recalibrated in the laboratory, the calibration factors were, $A = 5044 \pm 56$ cts/pg ($n = 8$) and $B = 5256 \pm 70$ cts/pg ($n = 8$), indicating a 20.6% and 16.8% decrease in response for the A and B cartridges, respectively, over the course of the field deployment. The data have been corrected for this drift by assuming the decrease in response was linear over instrument operation time. In support of this, we noted the absence of a temporal trend in our oceanic boundary layer data throughout the deployment, which remained near a constant value of 1.2 ng m$^{-3}$ and significant jumps in oceanic boundary layer GEM were not observed between flights. Instrument background measurements before and after each flight were always small and constant.

Instrument blanks were attempted in flight using “Zero Air” generated in situ by passing the sample air stream through a commercially prepared activated carbon canister (Resisorb, Tekran Inc.). Meaningful blanks were not obtained until it was discovered that a fraction of the mercury that was adsorbed on the activated carbon at low altitude desorbed at high altitudes. The data reported here are uncorrected for a gold cartridge-derived fluorescence background.

2.2. Mercury Measurement Issues Unique to Aircraft Operation

Previous work has shown the Tekran instrument responds to variations in pressure. This could come about through pressure effects on air sample volume, gold cartridge collection efficiency, carrier flow, and detector response. The sample airflow rate is controlled using a mass flowmeter and a circuit which regulates the sample pump’s speed to maintain a constant mass flow rate (1.50 slpm; $T = 273.15$ K, $P = 1013.25$ hPa). At higher altitudes (lower air pressures) the pump was found to be undersized and unable to maintain constant mass flow. Fortunately, a constant mass flow rate is not required as the mass flow rate from the meter is monitored and corrected for pressure effects.
integrated over the time of a sample to derive the total sample volume in standard liters (STP). It is this total sample volume, which is used to calculate GEM in ng m\(^{-3}\).

[14] Ebinghaus and Slemr [2000] demonstrated that the Tekran 2537A’s gold tube traps quantitatively collected GEM (100% efficiency) to their test limit of 500 hPa (~5.5 km altitude). Further, the Gormley and Kennedy [1949] equations indicate diffusional wall loss in a tube is independent of pressure under laminar constant mass flow. On the basis of these works, we assumed the diffusive collection of GEM on the collecting gold spheres is 100% up to our highest sample altitudes (~8 km or ~400 hPa). This was verified in the laboratory and the results are presented below.

[15] The Tekran’s detector system has been shown to be sensitive to operating pressure [Ebinghaus and Slemr, 2000; Friedli et al., 2003b]. These investigators determined and applied an empirical pressure-correction algorithm to their airborne observations. The pressure dependence of the Tekran’s cell was determined by us for pressures from 800 to 1026 hPa. The pressure response was linear and 0.118% per hPa (A) and 0.114% per hPa (B). These values were very similar to those reported by the above authors. To avoid the need for this pressure correction, an electronic pressure controller (MKS Series 640) was installed at the exit of the cold vapor atomic fluorescence detection cell to maintain a constant detection system operating pressure in the cell and in the gold cartridge during the thermal desorption process. A set point pressure of 1026 ± 2 hPa was selected to ensure the cell pressure always exceeded ambient pressure during the field experiment. The controller was tuned to avoid oscillations triggered by surges in Ar flow, which resulted from the Tekran’s internal flow switching sequences.

[16] Premission the MKS pressure control system and the assumption about the gold tube’s sampling efficiency were evaluated. A Tedlar bag containing a known concentration of GEM was prepared. The bag was plumbed to the Tekran inlet through a Teflon throttle valve, which allowed the inlet pressure to be varied. The indicated Hg concentration remained constant as the inlet pressure was reduced stepwise over the range expected to be experienced aboard the WP-3.

[17] Postmission the Tekran’s pressure dependence was reevaluated in the laboratory over a wider pressure range from 830 to 300 hPa, 100 hPa below the operational range experienced during the field experiment. Manual injections into sample flows at 600 and 300 hPa resulted in responses of 5864 and 6002 counts per picogram, indicating that the response was relatively unaffected (~2%) even at these lowest pressures.

[18] As will be presented below, the ITCT2K2 GEM altitude distributions showed, contrary to expectation, a significant decrease in concentration with altitude above 3 to 4 km. We searched for possible experimental or instrumental artifacts that could have contributed to decreasing GEM concentrations with increasing altitude. Plumbing leaks were one possibility. In the field, the manufacture’s leak check procedures were followed and no leaks were found. Postmission, a second set of checks were performed as part of our research into the collection efficiency of the gold cartridges and detector pressure response. Again, no leaks were found.

[19] One additional note about data reduction is needed. At about 6000 to 7000 m altitude and above, the output of the Tekran instrument reported “0” as the amount of GEM collected fell below the instruments internal criteria for “good data.” This resulted from two factors. First, the sampled air total volume fell below a prescribed minimum sample volume, as the pump had insufficient capacity at higher altitudes. Second, the small sample volume coupled with low mercury concentrations at these altitudes failed to deliver a sufficient quantity of mercury during desorption to generate a signal above a preset threshold value relative to baseline that the instrument could detect and integrate. Under either condition the instrument reports “0” GEM. This resulted in an effective high-altitude detection limit of 0.2 ng m\(^{-3}\). These “0” samples were removed during postmission data reduction. The removal of these “0” values introduced a potential positive bias in our observations at highest altitudes.

3. Results

[20] Approximately 1600 observations with quantifiable GEM concentrations were made on 11 WP-3 flights during ITCT2K2. The flight tracks are shown in Figure 2 for the 9 sorties based from Monterey, CA. Another 60 samples fell below practical detection limits, 0.1 ng m\(^{-3}\) from the surface to ~5 km and 0.2 ng m\(^{-3}\) above ~5 km. Each sample was collected over 150 s, with a sample’s start and stop times defining the GEM time base used to present the data. Most other instruments had faster response times and reported data at much shorter intervals, e.g., CO, H\(_2\)O, or O\(_3\). The faster measured constituents have been averaged over the GEM sample start-stop time and merged to the
GEM timelines for GEM, CO (solid circle), O₃ (open circle), NO₂, acetonitrile, SO₂, and the ratio of propane to ethane on 5 May 2002. All data lie between 4 and 8 km. This flight encountered three elevated GEM plumes (3, 4 and 0.7 ng m⁻³) at 23.0–23.25, 23.25–23.4, and 23.9–24.0 UT. Each GEM plume is associated with a variable mix of copollutants. CO, O₃, SO₂, and acetonitrile have been averaged (merged) to the GEM time base. Propane/ethane is from whole air samples and is plotted on an independent time base—midpoint of whole air sample collection time.

The data presentation is organized to highlight three case studies: a long-range transported (LRT) plume(s), the sampling of the Los Angeles Basin (LAB) air, and the sampling of an upper tropospheric–lower stratospheric (UTLS) impacted air mass. Figures 3–5 highlight the LRT plume encountered on 5 May 2002. Figures 6–8 depict the air composition over the Los Angeles Basin. Figures 9–11 show the influence of UTLS air on GEM. Figure 12 shows a summary of GEM measured in ITCT2K2 and is compared with data from ACE-Asia. Figure 13 shows, for completeness, the GEM vertical distribution measured for the other 8 flights.

CO, O₃, ethyne, ethane and propane are shown for all three case studies. Acetonitrile, SO₂, and NO₂ observations are shown for the LRT and LAB pollution cases. CO was chosen as a general long-lived pollution tracer and because the GEM-CO ratio was used previously to identify Asian plumes reaching North America [e.g., Jaffe et al., 2005]. O₃ served as both a long-lived pollution tracer as well as an indicator of stratospheric influence. NO₂, SO₂, ethyne, ethane and propane were selected as representative pollution tracers of moderate lifetime (1–10 days). In addition, both ethyne to CO and propane to ethane ratios have been used as qualitative indicators of air mass processing.

During the project, only 5 and 10 May exhibited large CO plumes (CO > 100 ppb over background) indicating lofted Asian air masses with rapid long-range transport to the coast of California. The 5 May plume (presented as the 10 May example) was complicated by mixing with stratospherically influenced air [Cooper et al., 2004], and was possibly further influenced by North American emissions as its parcel trajectory (not shown) showed it to have traveled at low altitudes along the coast.
3–4 km, above which it distinctly decreased with altitude. We have called this intermediate altitude the GEM mixed depth and it is listed for each flight in Table 1. A composite vertical distribution of GEM for all 11 ITCT2K2 flights is shown in Figure 12. Thereon the median, 25th and 75th percentiles are plotted for 1 km layers altitude. Note the pronounced decrease in GEM from the lower altitudes to the upper altitudes. The GEM decrease with altitude was unexpected and is commented upon more fully in the discussion section.

Table 2 lists the mean GEM, CO, O$_3$, and H$_2$O values for contiguous-in-time observations from each flight for the lowest altitudes sampled over the ocean. Mean latitude, longitude, and a land reference point are listed, too. The data were selected to have ambient pressures less than 985 hPa, which corresponds to a maximum sample altitude of 250 m. This thin vertical interval was chosen because a deeper 0–1 km layer often selected air mass samples from both above and below the MBL top. The GEM decrease with altitude was unexpected and is commented upon more fully in the discussion section.

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A comparison of the ITCT2K2 0–250 m data with other Pacific data is given in Table 3. The ITCT2K2 data were lower on average and exhibited a lower range of values than previous work, although our mean falls within the range of these studies. Table 3 also shows a comparison of a subset of the ITCT2K2 data with elevated
mountain sites, Cheeka Peak, WA, and Mt. Batchelor Observatory, OR. The ITCT2K2 data for this comparison were selected to be north of 40°N and between 2.3 and 3.3 km so as to best compare with MBO latitude and altitude. Selected MBO data from 15 April to 27 May were chosen to provide a ±10-day overlap with the ITCT2K2 sample period. The MBO data were from 2004 and 2005 as MBO was not operational in 2002. The ITCT2K2 GEM data were about 50% lower. Not listed are comparisons for CO, O₃, and water vapor, which showed the selected ITCT2K2 samples to be 50% lower in water vapor, 20% lower in CO and 30% higher in O₃. This indicated the ITCT2K2 air masses to be less polluted and to have originated at higher, colder and drier altitudes. To match MBO values, required using ITCT2K2 observations from 0 to 2 km and from south of 35°N.

[29] The altitude variation was the predominant feature in GEM distributions. To isolate that effect from other analyses, the data from all flights were sorted into 1 km layers or bins. GEM displayed a relatively constant background concentration in each layer and it was possible to visually identify GEM peaks or “plumes” larger than 0.15 ng m⁻³ above the local background. Similarly, CO, O₃ and a suite of other chemicals were examined for deviations above background, which would indicate a “plume” event. The threshold for O₃ and CO events was 5 and 10 ppb, respectively. Events having O₃, CO and GEM positively correlated were interpreted as pollution transport events. There were 2 clear events on 29 April and 10 May where CO and GEM were positively correlated, but with O₃ negatively correlated. It is our interpretation these samples corresponded with stratospheri-
GEM events have visually detectable higher CO nor did all CO events have visually detectable higher GEM.

3.3. Asian Plume Case

[30] The vertical and horizontal distributions of GEM were significantly altered by distant and local sources on a portion of only four flights, 5, 13, 17 and 19 May 2002. The flight on 13 May was the Los Angeles sortie and is presented below. The flight on 19 May was the transit flight back to Colorado and the large GEM values were encountered on the decent into Denver. The 17 May high GEM appears to have been in North American outflow. On 10 May plumes of other species were encountered but changes in GEM were small, on the order of 0.2–0.3 ng m\(^{-3}\), and were near stratospherically impacted air with GEM dips of \(-0.3\) to \(-0.4\) ng m\(^{-3}\) associated with 50 ppv spikes in O\(_3\). The 150 s GEM sample integration interval may have smeared the pollution and the UTLS air masses together during this portion of the flight. On 5 May, plumes were encountered at an altitude of 5–7 km and had GEM up to 4 ng m\(^{-3}\). This flight paralleled the coast extending from 33 to 44\(^\circ\)N (Figure 2). The plume was located to the north with clean air to the south.

[31] During the 5 May 2002, flight three distinct plumes were encountered based upon CO and GEM. The selected parameters are shown as a function of time on Figure 3 for the plume period. Figure 4 shows vertical distributions of selected parameters for the entire flight. A backward air trajectory for the plume period is shown on Figure 5. Plume “A” is defined to extend from 23.0 to 23.25 UT on Figure 3 and has GEM concentrations of 2.5–3.0 ng m\(^{-3}\). Plume “B” extended from 23.25 to 23.35 UT and has the highest GEM at 4.0 ng m\(^{-3}\). Plume “C” was less distinct in its...
GEM concentration, extended from 23.9 to 24.05 UT and had GEM <0.3 ng m\(^{-3}\) above background. CO mixing ratios in the A, B and C plumes were 125, 100 and 150 ppb, respectively, above a background mixing ratio of 125 ppb. \(\text{O}_3\) was relatively constant at 75 ppb. It can be seen that (1) plumes A and C contained elevated \(\text{NO}_2\); (2) plumes A, B, and C contained elevated acetonitrile; (3) plumes A, B, C had elevated \(\text{SO}_2\) but the highest \(\text{SO}_2\) was just prior to plume C; and (4) plumes A and C had the “freshest” air, propane to ethane \(\sim 0.35\), compared to B, \(\sim 0.25\).

This illustrated the variability in GEM and other species composition in a LRT plume(s) event. The \(\text{D}_{\text{GEM}}:\text{D}_{\text{CO}}\) ratios (ng m\(^{-3}\)/ppb) were 0.024, 0.035 and 0.002 in plumes A, B and C, respectively. Elevated ethyne \(\sim 500 \text{ ppt}\) and ethane \(\sim 2000 \text{ ppt}\) in the plume at altitude is shown on Figure 4.

Figure 4 facilitates identification of background air with which to compare plume concentrations. Data when GEM is greater than 1.4 ng m\(^{-3}\) are circled in blue. Data when \(\text{O}_3\) was less than 45 ppb are circled in red. The red circled points have the highest relative humidities, lowest CO and lowest GEM at the altitude of the plume. Separately, it was determined the lowest ethyne, ethane, and propane-ethane ratio \(\sim 0.06\) corresponded with the lowest \(\text{O}_3\) at plume altitude. This flight encountered some of the “cleanest” most processed air at its southern end as well as the Asian plume(s) to the north.
Backward trajectory analysis suggested plume A originated from near the marine boundary layer off the coast of China and reached the sampling site after ~10 days. The trajectory for plume B is shown in Figure 5 and indicates this air moved across the Pacific at about twice the speed of the lower one, and ~5 days before our observations, it passed over Mongolia at >8000 m. The source of the fast-moving plume was not known, but de Gouw et al. [2004] identified acetonitrile as one of the pollutants in these plumes, which suggested a biomass burning contribution and deep pyroconvection.

### 3.4. Los Angeles Basin Case Study

The 13 May flight within the Los Angeles Basin showed a broadly distributed urban GEM source region. The entire flight concentration timeline for GEM, CO, O$_3$, NO$_2$, acetonitrile, SO$_2$ and propane-ethane ratio are shown in Figure 6. Vertical distributions of GEM, CO, O$_3$, relative humidity, potential temperature, and propane-ethane ratio are shown on Figure 7. Data points when GEM was greater than 1.4 ng m$^{-3}$ are circled in blue. Data when O$_3$ was less than 45 ppb are circled in red. The depth of the urban boundary layer was ~3.5 km as noted in CO, relative humidity (rel-hum), and propane to ethane ratio. The highest GEM, 2.2 ng m$^{-3}$, was confined to lower altitudes and appeared capped at the main inversion: sharp potential temperature increase at 1–1.5 km. Highest CO, propane to ethane ratio, O$_3$, NO$_2$, and SO$_2$, were found at these lowest altitudes, too. There was a high correlation between primary pollutant species. This is illustrated for CO-GEM on Figure 8, which shows the correlation between these 2 species in the 0–1, 1–2, 2–3 and 3–4 km layers over the Basin. The CO-GEM slope varied from 0.0015 to 0.0024 with an average value of 0.0018 (ng m$^{-3}$ ppb$^{-1}$).

The low O$_3$ data (<45 ppb) were principally located along the Monterey to Los Angeles transit line at 3.5 km. This transit showed some of the highest relative humidity aloft, lowest CO and mixed propane-ethane ratios (19.5 to 20.5 UT on Figure 6) suggestive of air mass mixing along the coast.

### 3.5. Stratospheric Influence Case Study

GEM was observed in stratospherically influenced air masses on 29 April 2002 and 10 May 2002. The 29 April case provided a clearer demonstration of stratospheric influence with O$_3$ values approaching 300 ppb. Figure 9 shows vertical distributions of GEM, CO, O$_3$, relative humidity, potential temperature, and propane-ethane ratio. The UTLS air began at about 6 km and extended beyond the highest sample altitude (8 km). This was evident in potential temperature, relative humidity, GEM, CO, and most clearly in O$_3$. Ethyne and ethane also showed a decrease with increasing O$_3$ above ~80 ppbv.

The anticorrelation between O$_3$ and GEM is illustrated in Figure 10. The data were for the altitudes above 6 km. The data showed a GEM to O$_3$ slope of ~0.0038 ng m$^{-3}$ ppb$^{-1}$. Extrapolation of the line to GEM = 0 yielded an O$_3$ value of 333 ppb. There was also an anticorrelation between CO and O$_3$ and which is not shown. In the UTLS air, there was a

![Figure 10](image1.png)

**Figure 10.** GEM and O$_3$ correlation above 6 km in the stratospherically influenced air on 29 April 2002. The slope, intercept and regression coefficient are listed. Extrapolation of the line to GEM = 0 would yield an O$_3$ value of 333 ppb.

![Figure 11](image2.png)

**Figure 11.** Same as Figure 10 except for GEM and CO. The extrapolation of the line to GEM = 0 yields a CO value of 54 ppb.
very strong correlation between GEM and CO. This is shown on Figure 11. The GEM-CO slope was 0.0137 ng m$^{-3}$ ppb$^{-1}$, which we propose was driven by the relative loss of each going into the stratosphere. Extrapolating to 0 GEM yielded a CO value of 54 ppb. We have also fit the O$_3$-CO data for this period and the extrapolation of CO to 54 ppbv yielded an O$_3$ value of 333 ppb, a value below tropospheric mixing ratios and in keeping with UTLS values. The 3 variables GEM-CO-O$_3$ yielded a self-consistent end-member for lower stratospheric air mixing with upper tropospheric air of 0–54–340, respectively.

Figure 9 also showed that at the base of the layer there appeared to be a convected layer with high relative humidity, CO, GEM, ethane and propane to ethane ratio but with lower O$_3$. The vertical distribution of many species below the UTLS indicated this was one of the more vertically well-mixed sampling days. The flight on 24 April was similar in this regard (not shown). Both had nearly constant GEM with altitude, until UTLS encounter.

4. Discussion

Four aspects of the ITCT2K2 GEM data set were notable in their implications to mercury in the atmosphere and the interpretation of its natural atmospheric cycles. First is the vertical gradient in GEM. Second is the GEM observations implying a 100-day residence time for GEM. Third, the apparent quantitative loss of GEM in the UTLS is a new finding of undetermined significance to GEM cycling and may afford a new tracer for UTLS air exchange. Fourth, the GEM-CO ratios in Asian plumes, in urban air, and the UTLS vary by well over an order of magnitude. Last, new GEM questions are posed.

4.1. Vertical Gradients of GEM

The earliest GEM vertical sounding observations of Abramovsky et al. [1977] revealed a strong negative gradient of more than an order of magnitude from the surface to 4 km. Kvetkus [2003] also found the concentration of GEM to be largely invariant from the surface to 7 km, with concentrations falling 60% from the surface to 4 km. Ebinghaus and Slemr [2000] obtained two vertical soundings. One of these soundings showed no significant gradient and the second showed a strong negative gradient from the surface to 2.3 km and no GEM gradient from 2.3 km to 3.7 km. R. H. Ebinghaus (personal communication, 2006) also reported a negative gradient in the lowest 0.2 km from a coastal tower. In an extensive series of soundings over southeastern Canada, Ontario, and the Canadian Arctic, Banic et al. [2003] found the concentration of GEM to be largely invariant from the surface to 7 km, with concentrations of 1.5 to 1.7 ng m$^{-3}$. Landis [2003] in a summer field program over Florida observed GEM to have a negative gradient with altitude.

Over the eastern North Pacific in ITCT2K2 GEM was found to have a generally modest to strongly negative vertical gradient. Whole flight composite soundings show...
three different and rather distinctive characteristics. These are presented in Figures 4, 7, 9 and 13. The principal characteristic is a significant, generally negative, gradient, with the GEM concentration falling as much as half an order of magnitude between sea level and 7–8 km, as on the 2, 5, 6, 10, and 11 May flights. The second characteristic can be seen in a few cases, such as 5 May (Figure 4), where GEM had little vertical gradient to ~4 km, when it became negative, but with one or more GEM-rich layers. In these cases the layer was also coincident with a strong CO increase (see both Figures 3 and 4). The third characteristic is seen on 25 and 29 April and 5 May (Figures 4 and 13), where below 2 km, GEM showed little discernible vertical gradient. Above this mixed region, the gradient once again became negative.

[45] Surface sources, atmospheric stability, slow chemical oxidation in the troposphere and depleted GEM in UTLS impacted air may have all played a role in establishing the vertical gradient in GEM observed in ITCT2K2. GEM clearly has surface sources especially on land. The ocean can be a source regionally and may buffer GEM concentrations in the MBL. Photochemistry slowly removes GEM by reaction with radicals. The upper troposphere lower stratosphere was depleted in GEM and inferred stratospheric concentrations were 0 ng m$^{-3}$. The flight plans selected for meteorologically stable conditions and suppress mixing and

**Figure 13.** Flight by flight composite of GEM vertical distributions. Data from flights on 29 April, 5 May and 13 May 2002 were previously shown and not repeated here. GEM decreases slightly with altitude up to 3 to 4 km and then decreases at different rates to the highest sample altitudes. The change in GEM “lapse” rate at 3–4 km is paralleled in the potential temperature lapse rate ($\theta$ is not shown here but was previously shown in figures for the flights on 29 April, 5 May, and 13 May).

<table>
<thead>
<tr>
<th>Date</th>
<th>Stratospheric Influence (UTLS)</th>
<th>GEM Mixed Depth, m</th>
<th>GEM + CO Plumes$^a$</th>
</tr>
</thead>
<tbody>
<tr>
<td>25 Apr 2002</td>
<td>6–8 km</td>
<td>3000</td>
<td>H</td>
</tr>
<tr>
<td>29 Apr 2002</td>
<td>6–8 km</td>
<td>5000</td>
<td>H</td>
</tr>
<tr>
<td>2 May 2002</td>
<td>2500</td>
<td>H</td>
<td>H</td>
</tr>
<tr>
<td>5 May 2002</td>
<td>4000</td>
<td>H</td>
<td>H</td>
</tr>
<tr>
<td>6 May 2002</td>
<td>5000</td>
<td>H</td>
<td>H</td>
</tr>
<tr>
<td>10 May 2002</td>
<td>2–5 km</td>
<td>3500</td>
<td>LH</td>
</tr>
<tr>
<td>11 May 2002</td>
<td>3500</td>
<td>H</td>
<td>H</td>
</tr>
<tr>
<td>13 May 2002</td>
<td>2500</td>
<td>H</td>
<td>LH</td>
</tr>
<tr>
<td>15 May 2002</td>
<td>step at 1000, 4000</td>
<td>LH</td>
<td>LH</td>
</tr>
<tr>
<td>17 May 2002</td>
<td>step at 1000, 4000</td>
<td>LH</td>
<td>LH</td>
</tr>
<tr>
<td>19 May 2002</td>
<td>NA</td>
<td>NA</td>
<td>LH</td>
</tr>
</tbody>
</table>

$^a$L, below 4 km; H, above 4 km.
Table 2. Mercury, Ozone, Carbon Monoxide, and Water Vapor

<table>
<thead>
<tr>
<th>Flight</th>
<th>Samples</th>
<th>Latitude, °N</th>
<th>Longitude, °E</th>
<th>Hg, ng m⁻³</th>
<th>CO, ppb</th>
<th>O₃, ppb</th>
<th>H₂O, hPa</th>
<th>Location</th>
</tr>
</thead>
<tbody>
<tr>
<td>20020425</td>
<td>1</td>
<td>35.37</td>
<td>−126.97</td>
<td>1</td>
<td>123</td>
<td>37</td>
<td>11.99</td>
<td>~500 km W Morro Bay, CA</td>
</tr>
<tr>
<td>20020425</td>
<td>3</td>
<td>36.96</td>
<td>−123.04</td>
<td>1.1</td>
<td>138</td>
<td>40</td>
<td>10.85</td>
<td>~100 km W Monterey, CA</td>
</tr>
<tr>
<td>20020429</td>
<td>9</td>
<td>36.99</td>
<td>−127.17</td>
<td>1.1</td>
<td>150</td>
<td>46</td>
<td>8.86</td>
<td>~400 km W of Monterey, CA</td>
</tr>
<tr>
<td>20020429</td>
<td>1</td>
<td>37.00</td>
<td>−122.76</td>
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<td>148</td>
<td>50</td>
<td>8.20</td>
<td>~100 km W Monterey, CA</td>
</tr>
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<td>20020502</td>
<td>32</td>
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<td>−126.85</td>
<td>1.1</td>
<td>131</td>
<td>41</td>
<td>9.76</td>
<td>~100 km W of Trinidad Head, CA</td>
</tr>
<tr>
<td>20020505</td>
<td>1</td>
<td>41.04</td>
<td>−124.89</td>
<td>1.1</td>
<td>119</td>
<td>38</td>
<td>9.67</td>
<td>~100 km W of Trinidad Head, CA</td>
</tr>
<tr>
<td>20020505</td>
<td>1</td>
<td>36.85</td>
<td>−122.82</td>
<td>1.1</td>
<td>122</td>
<td>42</td>
<td>11.10</td>
<td>~100 km W Monterey, CA</td>
</tr>
<tr>
<td>20020510</td>
<td>2</td>
<td>41.02</td>
<td>−127.20</td>
<td>1.1</td>
<td>126</td>
<td>38</td>
<td>7.94</td>
<td>~300 km W of Trinidad Head, CA</td>
</tr>
<tr>
<td>20020511</td>
<td>5</td>
<td>47.75</td>
<td>−124.92</td>
<td>1.1</td>
<td>123</td>
<td>43</td>
<td>5.37</td>
<td>~100 km W of Aberdeen, WA</td>
</tr>
<tr>
<td>20020515</td>
<td>5</td>
<td>32.31</td>
<td>−128.54</td>
<td>1.2</td>
<td>129</td>
<td>48</td>
<td>10.85</td>
<td>~1100 km W San Diego, CA</td>
</tr>
<tr>
<td>20020515</td>
<td>6</td>
<td>37.42</td>
<td>−126.67</td>
<td>1.2</td>
<td>132</td>
<td>43</td>
<td>9.76</td>
<td>~400 km W of Half Moon Bay, CA</td>
</tr>
<tr>
<td>20020517</td>
<td>7</td>
<td>36.40</td>
<td>−122.56</td>
<td>1.3</td>
<td>118</td>
<td>36</td>
<td>11.82</td>
<td>~100 km W Monterey, CA</td>
</tr>
<tr>
<td>20020517</td>
<td>3</td>
<td>35.14</td>
<td>−124.20</td>
<td>1.2</td>
<td>112</td>
<td>34</td>
<td>12.31</td>
<td>~300 km W Morro Bay, CA</td>
</tr>
<tr>
<td>20020517</td>
<td>7</td>
<td>30.73</td>
<td>−129.32</td>
<td>1.2</td>
<td>106</td>
<td>27</td>
<td>13.37</td>
<td>~1900 km W of Baja, Mexico</td>
</tr>
<tr>
<td>20020517</td>
<td>6</td>
<td>34.67</td>
<td>−125.27</td>
<td>1.3</td>
<td>107</td>
<td>32</td>
<td>13.41</td>
<td>~500 km W of Lompoc, CA</td>
</tr>
</tbody>
</table>

Mean | 1.2 | 125 | 40 | 10.35

Std-Dev | 0.1 | 13 | 6 | 2.19

Table 3. GEM Comparison With Other Pacific Observations

<table>
<thead>
<tr>
<th>Site</th>
<th>Period</th>
<th>Mean (Median)</th>
<th>Range</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>Coastal East Pacific, &lt;250 m</td>
<td>27 Apr to 19 May 2002</td>
<td>1.2(1.2)</td>
<td>1.0–1.4</td>
<td>this work</td>
</tr>
<tr>
<td>Okinawa, 0 km</td>
<td>1 Jan to 31 May 2005</td>
<td>2.0(2.0)</td>
<td>1.4–4.7</td>
<td>Jaffe et al. [2003]</td>
</tr>
<tr>
<td>Mid N. Pacific, 0 km</td>
<td>1980–1986</td>
<td>~1.6</td>
<td>0.8–2.3</td>
<td>Fitzgerald [1995]</td>
</tr>
<tr>
<td>ACE-Asia, 0–1 km</td>
<td>Spring 2000</td>
<td>na(2.0)</td>
<td>1.7–2.3</td>
<td>Friedli et al. [2004]</td>
</tr>
<tr>
<td>Coastal East Pacific, 2.3–3.3 km north of 40°N</td>
<td>0.9 ~ 0.1</td>
<td>na(2.0)</td>
<td>1.7–2.3</td>
<td>this work</td>
</tr>
<tr>
<td>Cheeka Peak, WA, 1 km</td>
<td>Spring</td>
<td>1.5(na)</td>
<td></td>
<td>Weiss-Penzias et al. [2003]</td>
</tr>
<tr>
<td>Cheeka Peak, WA, 1 km</td>
<td>Summer</td>
<td>1.6(na)</td>
<td></td>
<td>Weiss-Penzias et al. [2003]</td>
</tr>
<tr>
<td>Cheeka Peak, WA, 1 km</td>
<td>Fall</td>
<td>1.5</td>
<td></td>
<td>Weiss-Penzias et al. [2003]</td>
</tr>
<tr>
<td>Washington coast, 0–3 km</td>
<td>Summer 2006</td>
<td>1.4(na)</td>
<td>1.1–2.0</td>
<td>authors unpublished work</td>
</tr>
<tr>
<td>Mt Bachelor, OR, 2.8 km</td>
<td>15 Apr to 27 May 2004</td>
<td>1.8(1.8)</td>
<td>1.5–2.5</td>
<td>Weiss-Penzias et al. [2006]</td>
</tr>
<tr>
<td>Mt Bachelor, OR, 2.8 km</td>
<td>15 Apr to 27 May 2005</td>
<td>1.7(1.7)</td>
<td>1.2–2.1</td>
<td>D. Jaffe (personal communication, 2006)</td>
</tr>
<tr>
<td>Canada, 0–7 km</td>
<td>Winter</td>
<td>1.1</td>
<td>0.4–1.7</td>
<td>Banic et al. [2003]</td>
</tr>
<tr>
<td>Ontario, 0–7 km</td>
<td>Summer</td>
<td>1.7</td>
<td>1.0–2.4</td>
<td>Banic et al. [2003]</td>
</tr>
<tr>
<td>Southeast Canada, 0–7 km</td>
<td>Spring</td>
<td>1.5</td>
<td>1.0–2.5</td>
<td>Banic et al. [2003]</td>
</tr>
</tbody>
</table>

The replenishment of GEM to the middle and upper troposphere except in the cases where distant convection has lofted plumes of GEM. These factors: sources on the surface, sinks at the tropopause and suppressed mixing, would all favor a negative gradient in GEM.

[46] In summary, the limited number of GEM soundings here and in the cited literature need be considered in conflict, as their differences in altitude structure may be entirely due to differences in the time since the last overturning of the troposphere.
4.2. UTLS GEM

[47] The stratospherically influenced air samples were a mixture of upper tropospheric and lower stratospheric air, UTLS. In the upper troposphere, concentrations of GEM were 0.3–0.5 ng/m$^3$. From the O$_3$-GEM regression analysis GEM was expected to be zero in UTLS air containing /C24 350 ppb O$_3$. The chemical mechanism of GEM loss in the UTLS is unknown. Holmes et al. [2006] have shown through model simulations that Br chemistry will deplete GEM in the UTLS.

[48] The quantitative conversion of GEM to, most likely, particulate mercury through RGM is supported by the observations of Murphy et al. [1998, 2003] and Ivlev et al. [1989]. They observed anomalous aerosol layers containing pHg in the UTLS region near the tropopause. The mechanism of stratospheric depletion and pHg production is uncertain.

4.3. CO to GEM Ratios

[49] The GEM to CO correlation or $\Delta$GEM/$\Delta$CO was used to identify Asian air masses impacting North America [e.g., Jaffe et al., 2005]. The key values were an Asian ratio observed in two plumes at Mt. Batchelor Observatory, ~0.0045 ng m$^{-3}$ ppb$^{-1}$ and a corresponding value on Okinawa, ~0.0056 ng m$^{-3}$ ppb$^{-1}$. In ITCT2K2 there was a wide span in visually deduced $\Delta$GEM/$\Delta$CO ratios ranging from 0.0000 to 0.035. The 5 May plume study yield values of 0.002, 0.024 and 0.035 ng m$^{-3}$ ppb$^{-1}$. The Los Angeles basin $\Delta$GEM/$\Delta$CO ratio was 0.002 ng m$^{-3}$ ppb$^{-1}$. The visual GEM event selection process failed to have a single occurrence of a positive GEM event without a zero or positive CO event, although there were a couple of CO events without a concurrent increase in GEM. Figure 14 shows the median event based $\Delta$GEM to $\Delta$CO ratio for each 1 km interval (vertically hatched bars). Each level had at least 3 events. Also plotted on Figure 14 is the $\Delta$GEM/$\Delta$CO slope determined from linear regression applied to CO-GEM data for each layer (diagonally hatched bar). Last, the heavy vertical dotted lines of Figure 14 indicate summary ratios from this work and the literature. A cursory reexamination of the ACE-Asia data for CO and GEM showed $\Delta$GEM/$\Delta$CO ranged from <0.001 to >0.02 ng m$^{-3}$ ppb$^{-1}$ in that program. In ITCT2K2 there is a trend for the $\Delta$GEM:$\Delta$CO ratio to increase with altitude. The vertical lines also suggest an increase in $\Delta$GEM:$\Delta$CO ratio from North American local to Asian transported plumes. The trend with height may then reflect a shift in GEM-CO source regions from western North American sources at low altitude to Asian or further upstream sources at high altitude.

4.4. GEM Lifetime Estimate

[50] The current estimated GEM residence time is ~1 year, where the residence time of GEM is defined by dividing the atmospheric burden by either the source rate or sink rate of GEM. A 1 year residence time was considered to be inconsistent with a pronounced negative vertical gradient observed in GEM during ITCT2K2. The GEM mean to variance ratio of all quantifiable samples was comparable to that observed in CO which has a photochemical lifetime on the order of 100 days. Such comparisons between a compound with unknown lifetime and compounds with known
Junge [1974] reasoned that the residence time of the residence time of the Std Deviation calculated tropospheric residence time for GEM by /C24 0 1 mean concentration Fitzgerald et al. in years, was given by:

\[ \Gamma \text{ (years)} = 0.14(\text{Mean concentration}) / \text{Std Deviation} \] (1)

[51] Junge [1974] reasoned that the residence time of trace gases should be inversely proportional to their variability and reasoned on theoretical grounds the residence time, \( \Gamma \) in years, was given by:

\[ \Gamma \text{ (years)} = 0.14(\text{Mean concentration}) / \text{Std Deviation} \] (1)

[52] He emphasized this approach was quantitatively valid (1) only “if the distribution of sources and sinks are similar” and (2) the trace gas of interest should be nearly at planetary steady state.

[53] GEM substantially meets these requirements. In ITCT2K2 cross Pacific transport at least partially homogenized the impact of strong Asian sources. In 11 flights directed to the study of intercontinental transport only two contained well defined plumes of GEM and other tracer compounds and these were sampled for a small fraction of the total sampling times. There were numerous small GEM events with highly variable GEM:CO ratios which suggested a blending of source regions. The putative chemical sinks of GEM: O3, HO, NO3, halogen and halogen monoxide radicals, were secondary oxidants whose natural and anthropogenic precursors shared in large part the same geographic source regions as GEM. Prior long-term studies have shown atmospheric concentrations of GEM are without annual trends, except after major volcanic eruptions [e.g., Fitzgerald et al., 1991; Ebinghaus et al., 2002; Weiss-Penzias et al., 2006] and suggested GEM was in global steady state.

[54] The ITCT2K2 data set was used to determine \( \Gamma \) for each flight using equation (1). These residence times are given in Table 4. The estimated residence times are seen to span \( \sim 50 \) to \( \sim 130 \) days and to average \( \sim 94 \) days. The GEM residence time estimate for the 5 May flight was 52 days, which was decidedly shorter than the estimates for the other flights. This was the day when LRT of an Asian air mass was observed at many levels and which had the highest GEM standard deviation. It is conceivable this flight was subject to strong source variations and did not meet the 1st criteria. In contrast, the 2 May flight did not have any identifiable plumes at any level (Table 1) and its residence time estimate, 92 days, was very near the mean value estimated above. For comparison, Fitzgerald [1986] using measurements made on a small Pacific island derived a similar residence time of 100 days using the Junge approach. Weiss-Penzias et al. [2003], reported a somewhat longer Junge residence time of 7 months in the summer at the Cheeka Peak Observatory, WA.

[55] In review, it was pointed out that one last caveat is in order. The Junge analysis above assumed the ITCT2K2 samples were meteorologically unbiased and the samples were collected over sufficient spatial-temporal scales for ergodicity to apply.

4.5. Study Implications

[56] The observed vertical gradient in GEM concentration if extrapolated over the world’s oceans, would lead to a one-third reduction in the estimated tropospheric GEM reservoir to 4000 tonnes. A reduction of the atmospheric burden leads to a concurrent decrease in residence time by one-third assuming sources and sinks remain constant.

[57] Separately, a reduction of GEM’s tropospheric residence time by about two thirds to the 100 day estimate made here required a reassessment of the sum of the source terms. Assuming current anthropogenic GEM estimates were correct at 3000 tonnes/yr, then the shorter residence time and atmospheric burden implied the estimated natural sources of GEM must increase by nearly a factor of 3 to about 9,000 tonnes yr\(^{-1}\). Conversely, assuming the natural sources were correct and total 3000 tonnes then a threefold increase in anthropogenic sources is needed. Either way the implied changes to the GEM budget would be large.

[58] The complexity revealed in the ITCT2K2 GEM measurements underscores the need for additional tropospheric observations of GEM. They were in considerable contrast with the more continental perspective provided by surface observation networks and earlier airborne observations.

[59] The budgetary points above were necessarily speculative and driven by the present study’s limited data set. They need to be confirmed or refuted directly by observations in the stratosphere, over the southern oceans and by more work in the Northern Hemisphere’s troposphere.

5. Conclusions

[60] GEM was observed over the eastern North Pacific in spring 2002 as part of the ITCT2K2 mission. The measurements were at the low end of the range of prior observations. GEM was found to be distributed in distinct layers and often together with chemical signatures of distant anthropogenic and natural sources. GEM concentrations in these few plumes, when contrasted with ACE-Asia’s observations near the Asian coastline, appeared to be nearly conserved during trans Pacific transport. GEM was generally correlated with CO and O3. Visually identified individual plume events showed variable changes in GEM relative to changes in CO. The highest concentrations of GEM were associated with combustion, high CO, and intermediate values of O3 (pollution and biomass fire). The lowest GEM concentrations were associated with UTLS impacted air high in O3, low in CO and very low in GEM. GEM was also observed to be low in air with low propane to ethane ratios indicative of highly processed (well aged) continental
air masses. The upper troposphere lower stratospheric air depletion in GEM suggested a heretofore unrecognized GEM sink. Also, the GEM-O$_3$ trend in the UTLS provided an additional quantitative measure of stratosphere-troposphere mixing. The combination of GEM’s observed variability, negative vertical gradient, and Junge residence time analysis suggested a GEM residence time nearer 100 days than the current value of 1 year.

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References


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